
The Fermilab Research Program

Presentation at
High-Energy Physics Advisory Panel
July 13, 2001
M. Shaevitz

Research Program Overview

- Tevatron Collider Run 2
 - Discovery potential Higgs, SUSY, Extra Dim.....
 - Precision measurements top, W, B-physics
- Neutrino Oscillation Program
 - MiniBooNE: Key experiment in sorting out the osc. situation
 - NuMI/MINOS: Definitive measurements in the atmospheric osc. region
- LHC Physics with CMS Detector
- Strong flavor physics program into the next decade with BTeV and Kaons
 - Along with possible next generation Neutrinos and other MI120/Low Energy Fixed Target program

Run 2 Program

(Initiated on March 1, 2001)

- Upgraded Fermilab accelerator complex
Goal is to increase number of available antiprotons/sec
 - New higher intensity 120 GeV Main Injector accelerator
 - New 8 GeV Recycler Ring made of permanent magnets
 - Improved antiproton cooling with both stochastic and electron cooling
 - Time between collisions will decrease from 396 nsec to 132 nsec
- Upgraded D0 and CDF detectors
 - Detectors, trigger, and data acquisition need to handle higher rates
 - Better micro-vertex, silicon tracker systems plus magnetic tracking
 - Detecting b-quark particles key to top, Higgs, and other physics
 - Better “hermetic” coverage ⇒ *“Missing energy” signatures*
- Extended runs from now through 2007-8 with continuous intensity upgrades

Run 2 Large Improvement over Run 1

- Run 2 Luminosity (integrated intensity) Plan
 - Run 2A: $\int L dt \sim 2 \text{ fb}^{-1}$ by end of 2003
 - Run 2B: $\int L dt \sim 15 \text{ fb}^{-1}$ by end of 2007
 - all at a cm energy of $\sim 1.96 \text{ TeV}$
- Compare this to Run 1 data with
 - $\int L dt$ of $\sim 0.10 \text{ fb}^{-1}$ ***P*** *All of Run 1 data*
 at cm energy = 1.80 TeV *= 1 week of running in Run 2B*

Process $p \bar{p} \rightarrow X Y$	Mass $Y \text{ GeV}/c^2$	Production sensitivity increase
$t \bar{t}$	175	1.4×150
$W H$	120	1.2×150
$\tilde{q} \tilde{q}$	300	1.8×150
$\tilde{g} \tilde{g}$	300	2.1×150

Sensitivity increase of Run 2 vs Run 1 200 - 300

Run 2 Physics

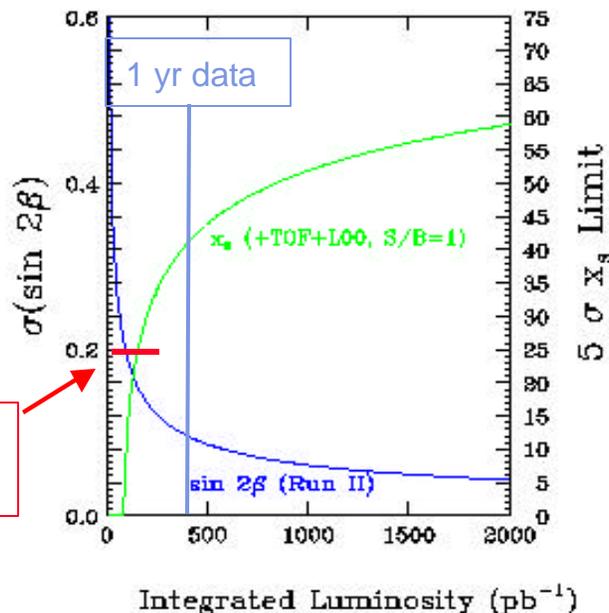
- **Run 2A** will explore at levels of roughly 40 - 80 times the sensitivity of Run I

- B physics: $\delta(\sin 2\beta) \sim \pm 0.05$
 x_s sensitivity to ~ 60
- Top physics: ~ 5000 top events
- Electroweak phenomena:
 $\delta(M_W) \sim \pm 50$ MeV
- QCD tests: Probe distance scales below 1 milli fermi
- Searches: new particle and extra dimensions

New g-2 Result

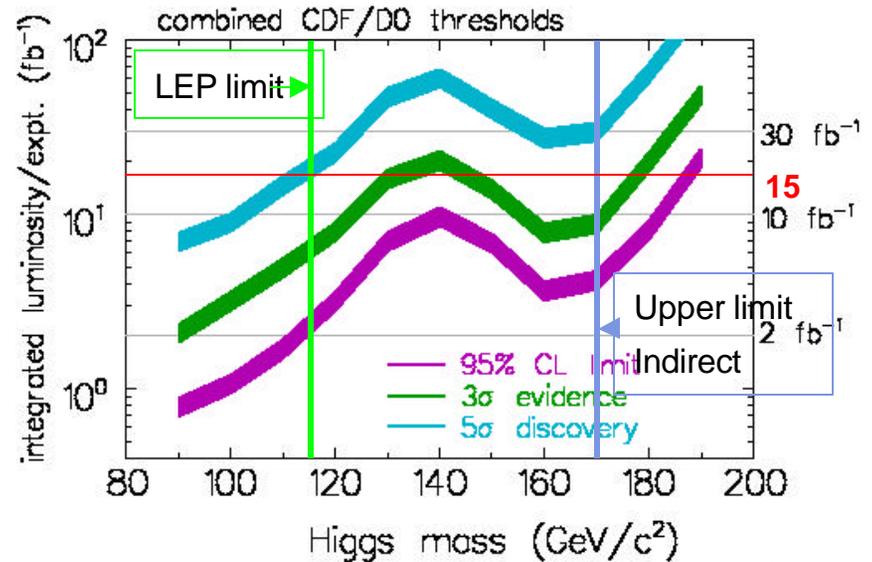
- Is SUSY just around the corner?

New BaBar/Belle result



- **Physics Potential in Run 2B** > 400 that in Run 1

- Every factor of two will open up new discovery possibilities
- Excellent discovery potential
 - Light Higgs
 - Entering farther into the domain of SUSY, extra dimensions



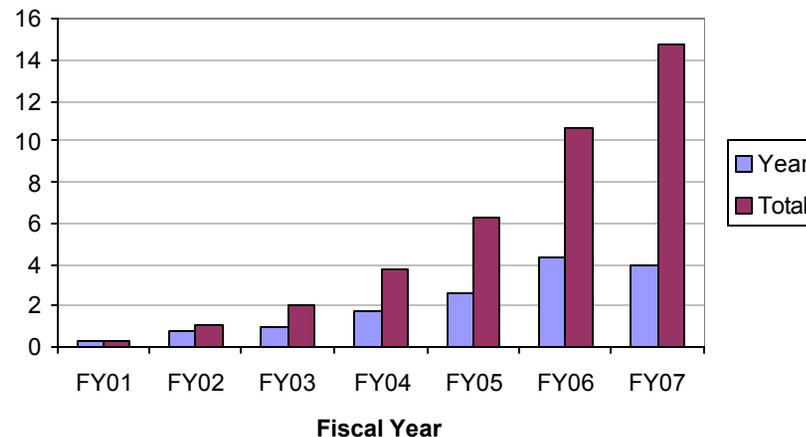
Run 2 Status

- Both CDF and D0 have rolled onto beamline
⇒ Run II started on March 1
- Initial machine and detector tune-up and commissioning
⇒ During March and April
- Collisions for physics during last two weeks of June
⇒ CDF $\sim 300 \text{ nb}^{-1}$ on tape and $> 1 \times 10^{30} \text{ cm}^{-2} \text{ s}^{-1}$ with
luminosity steadily increasing
- Goal is to reach 4 to 5×10^{31} luminosity by January
⇒ Collect integrated luminosity of 200 - 300 pb^{-1} by July, 2002

Run 2B Accelerator Upgrade Plans

Goal is to accumulate 15 fb^{-1} by 2007

- Key is to increase the number of antiprotons by x2-3
 - More protons on the antiproton target (x1.8)
 - Slip stacking plus brighter booster
 - Better antiproton collection efficiency
 - Upgrade Li-lens (\sim x1.5) and increase AP2-debuncher aperture (\sim x1.5)
 - Better Cooling and antiproton transfer efficiency
 - Including electron cooling in the recycler



Plans for Run 2B Detector Upgrades

- Main problem is radiation damage to silicon trackers and electronics at the $\sim 5 \text{ fb}^{-1}$ level
 - Upgrade strategy
 - Short shut-down in \sim summer, 2004 for installation (~ 6 months)
 - Replace silicon detectors and electronics with rad hard units
 - Improved triggering
 - Some other upgrades including computing
 - D0 and CDF upgrade projects being set up now
 - Technical Design Reports in Fall, 2001 \Rightarrow Baseline projects by Jan., 2002
 - Start on R&D projects
 - Joint project between CDF and D0 to develop new readout SVX4 chip
 - Begin R&D on sensors
 - Need both detectors to cover the physics
 - \Rightarrow Effectively doubles the luminosity

Enormous potential and major discoveries are likely
but we need to reach the Run IIb luminosity goals

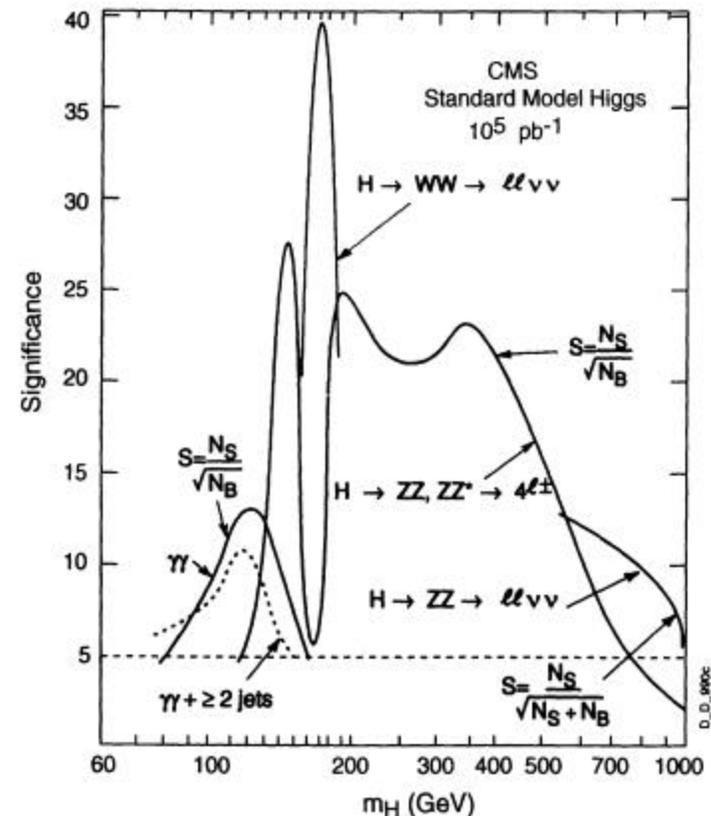
The LHC at Fermilab

Fermilab has a central role in two parts of the US effort on the Large Hadron Collider being constructed at CERN

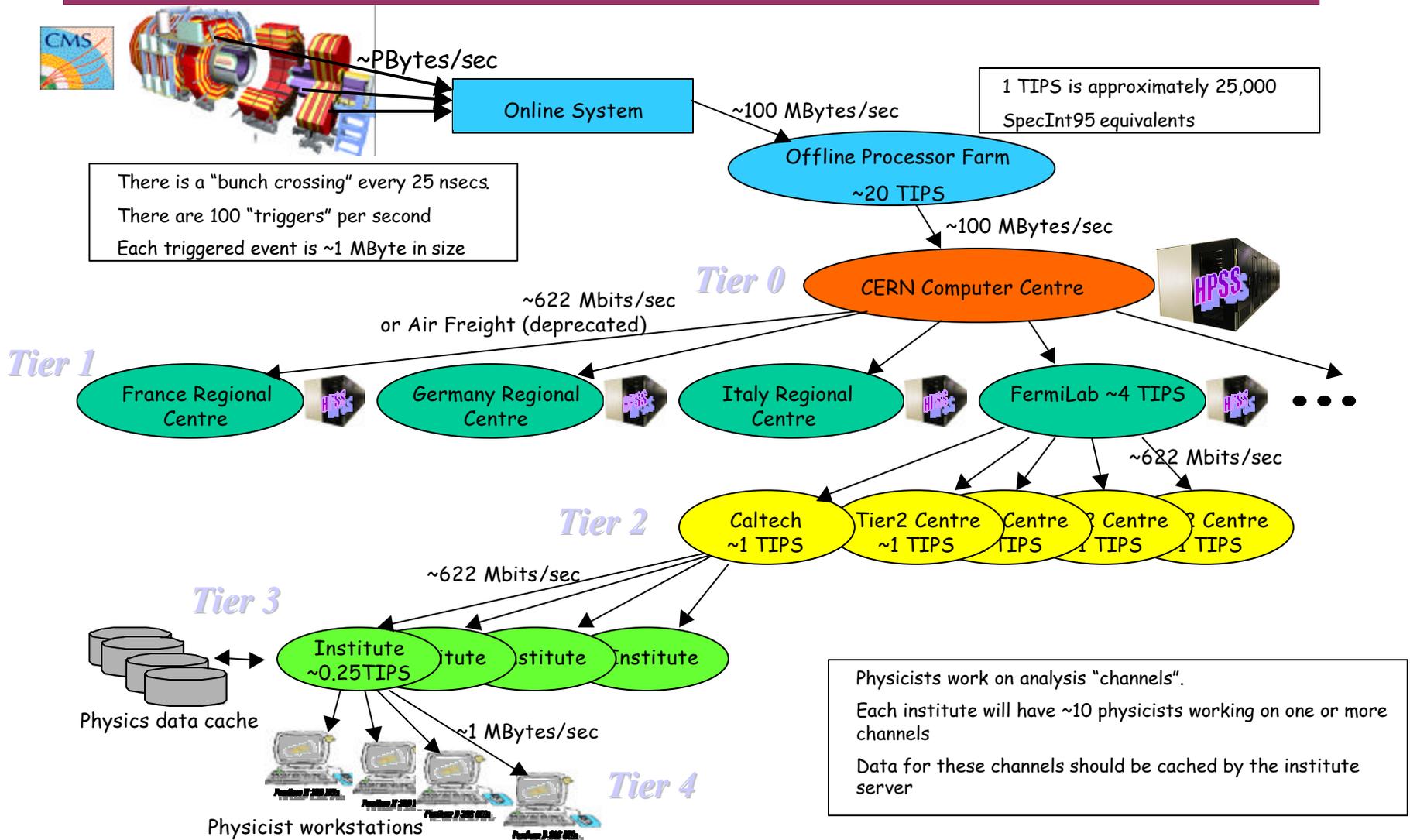
- US LHC Accelerator Project
 - The Project Manager Jim Strait manages a 3-laboratory collaboration building components for the LHC.
- US CMS Project
 - Fermilab is the host lab for the US CMS project.
 - We also are setting up the Tier I center for US CMS computing.

Fermilab participation allows US physicists to take a major role in building the LHC accelerator, building the CMS detector, and doing the science.

- Tevatron Run IIb will go up to ~ 180 GeV
- CMS will cover the range up to ~ 1 TeV: Mainly with $ZZ \rightarrow 4l$



CMS Software and Computing Project at Fermilab





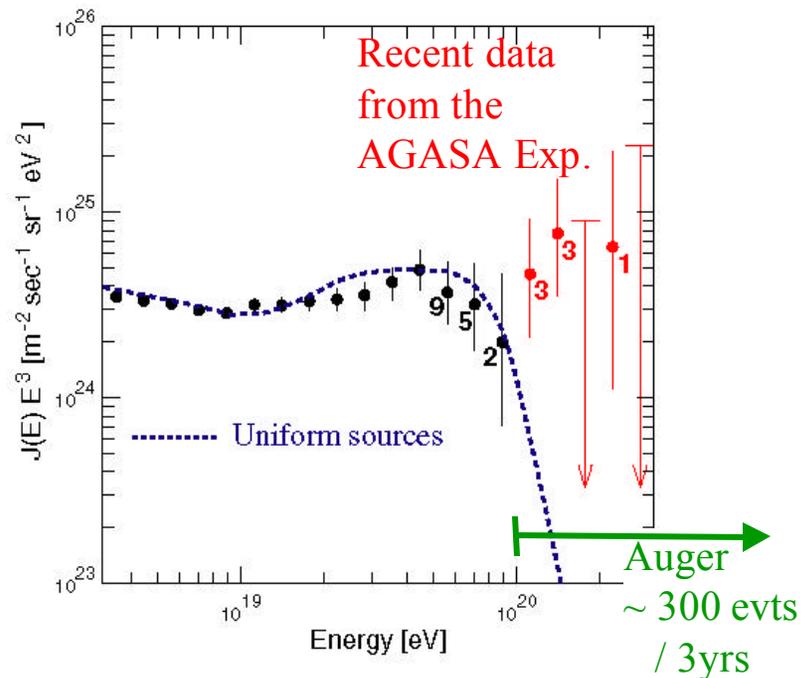
PIERRE
AUGER
OBSERVATORY

Astrophysics at Fermilab: CDMS, Auger, and SDSS

- Pierre Auger Project

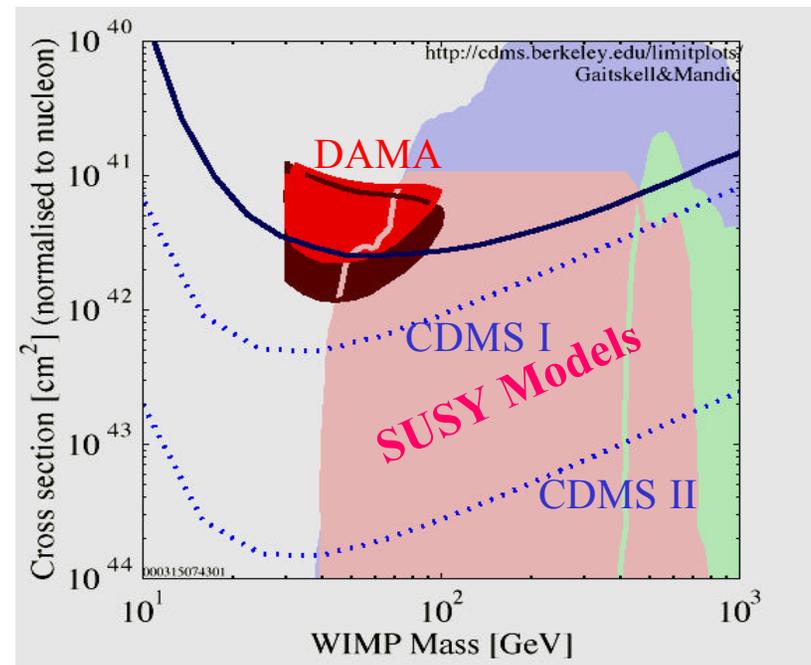
A study of cosmic rays 10^{19} eV - 10^{21} eV
Using very Large Ground Array in Argentina

- First fluorescence detector complete and operating → 1st events recorded
- All 40 of the “Engineering Array” deployed.
- Construction of the full observatory 2002 -2004.



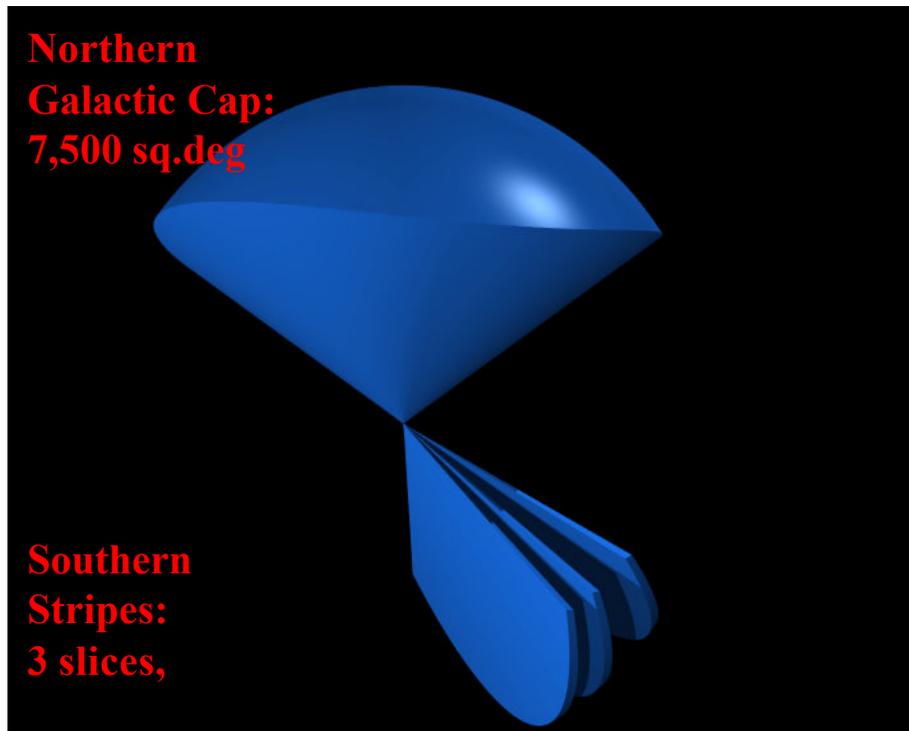
- CDMS (Cold Dark Matter/Wimp Search)

- Silicon and germanium crystal detectors at cryogenic temperatures (15 mK)
- CDMS I at Stanford in Summer, 2001
- CDMS II at Soudan, Minn.(with MINOS) to begin data early in 2002
 - Potential sensitivity 100 times better than current

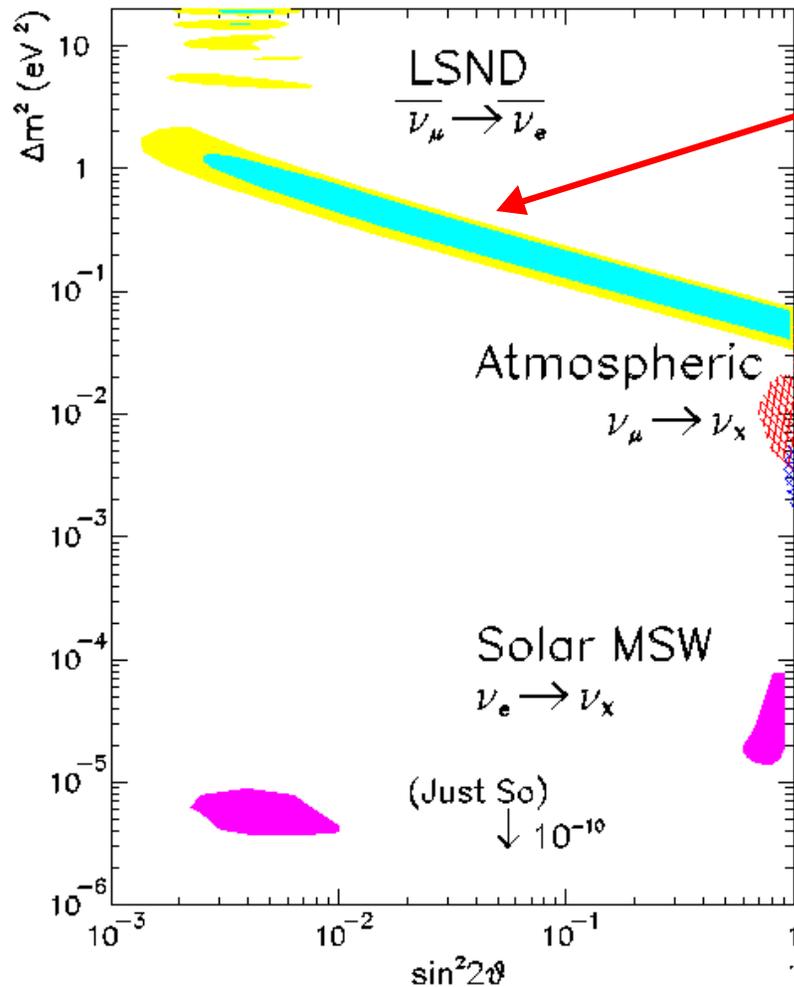


Sloan Digital Sky Survey (SDSS)

- Mapping the universe
 - Large imaging and spectroscopy survey
 - Covering 8000 sq.deg. or about 1/4 of sky
 - Digitize 100 million objects and spectra for million galaxies
- Science Accomplishments
 - Most distant and largest sample of **quasars**.
 - Large sample of **cool stars and brown (methane) dwarfs**.
 - Observation of **weak lensing**
 - This will yield a value for the Ω_m .
 - Large and uniform sample of **galaxy redshifts** gives knowledge of the galaxy luminosity function and properties



Fermilab ν Oscillations Program



- **MiniBooNE experiment:**

- Use neutrino beam from booster to confirm/refute positive LSND signal
- 1 GeV ν beam from 8 GeV booster to detector at 500m
- MiniBooNE crucial in understanding ν -osc. pheno.

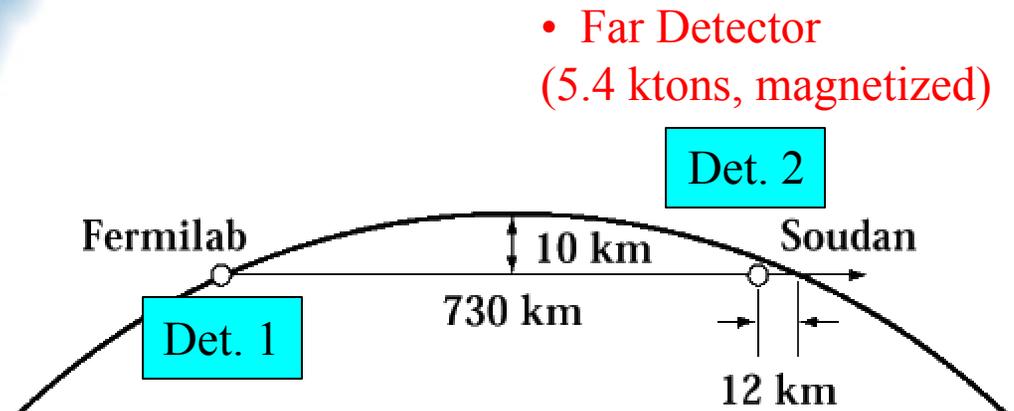
- **NuMI/Minos experiment:**

- Study atmospheric neutrino anomaly with terrestrial beam
- 3 to 20 GeV ν beam from 120 GeV Main Injector to detector at 738 km (Minnesota)
- Competition from Japanese K2K experiment and CERN to Gran Sasso (Italy)

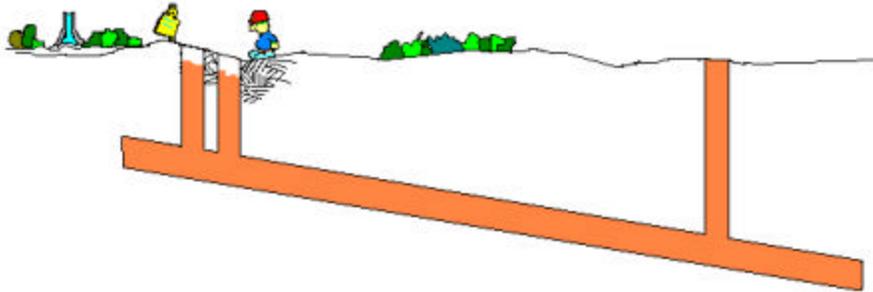
The NuMI/MINOS Experiment



Two Detector Neutrino Oscillation Experiment



NuMI/Minos Project

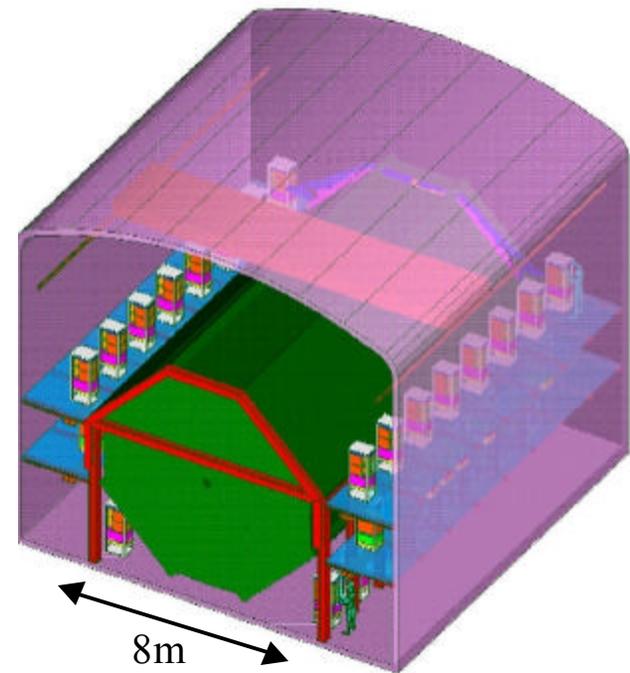


- Start of Soudan (Far Site) Construction 6/99
- Start of Fermilab Construction 3/00
- Start Far Detector Installation 8/01
- Start Near Detector Installation 6/03
- First Beam to Soudan 10/04

At Fermilab

- Submitting a rebaseline for the Project
 - Very difficult construction project at Fermilab
 - Cost and schedule under considerable stress
 - New baseline
 - Restoration of reasonable contingency: add 12 months (Oct.,04) and ~25% to TPC

(Official DOE CD-4 Start Operations is Sept.05)



At Soudan, Minnesota

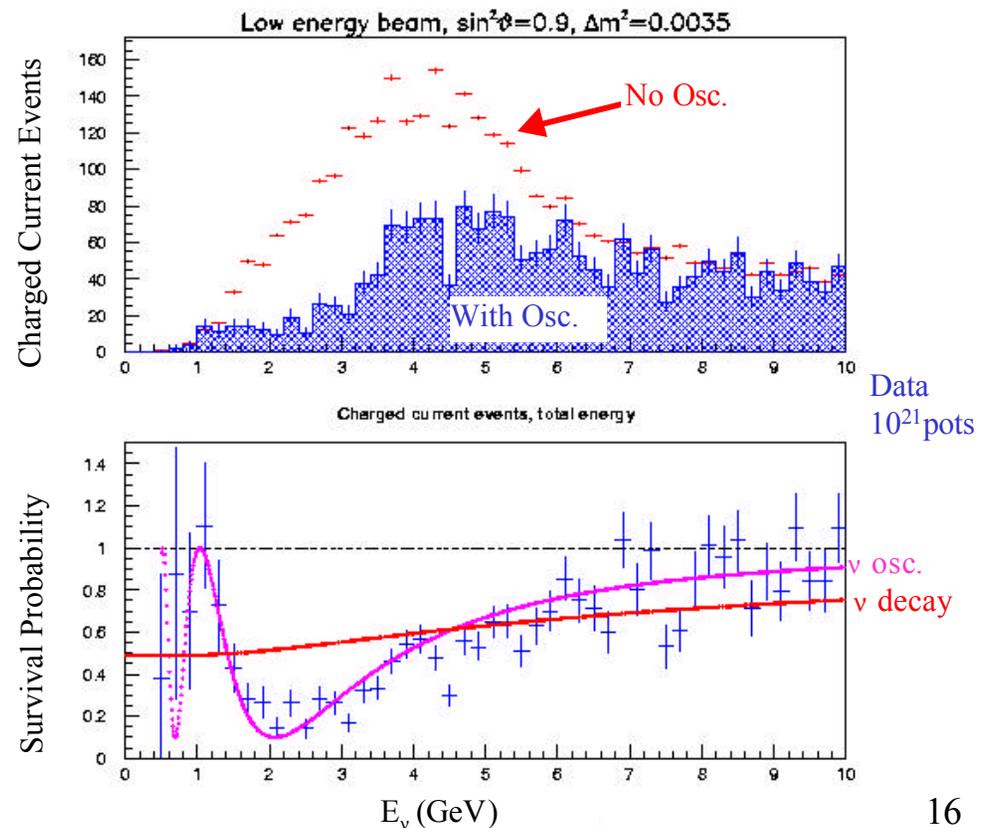
Oscillation effects observable at MINOS

Dm^2 sensitivity ($\sim 8 \times 10^{20}$ pot)

- CC energy analysis at low energy beam
- Includes Statistical & systematic errors

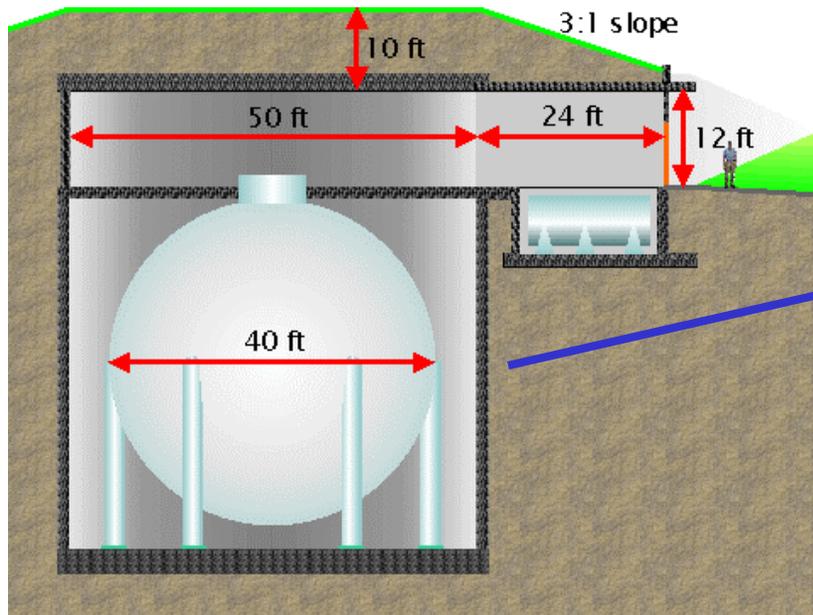
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Preview:
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See oscillatory behavior in
charged current rate vs E_ν

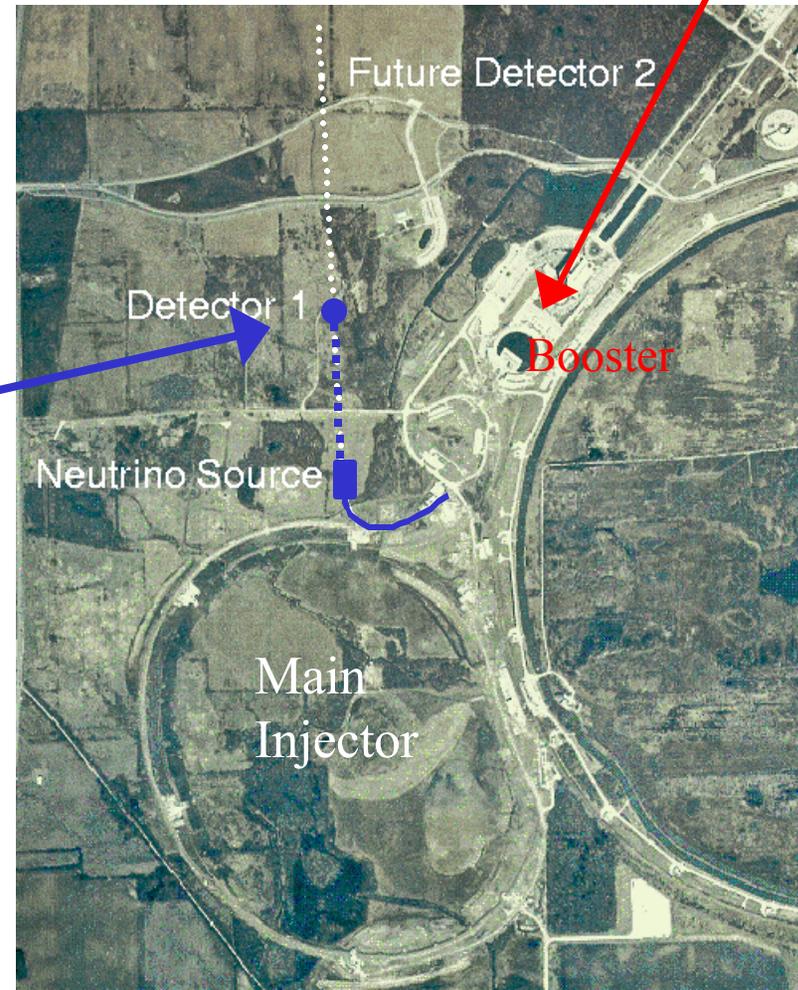


MiniBooNE Experiment

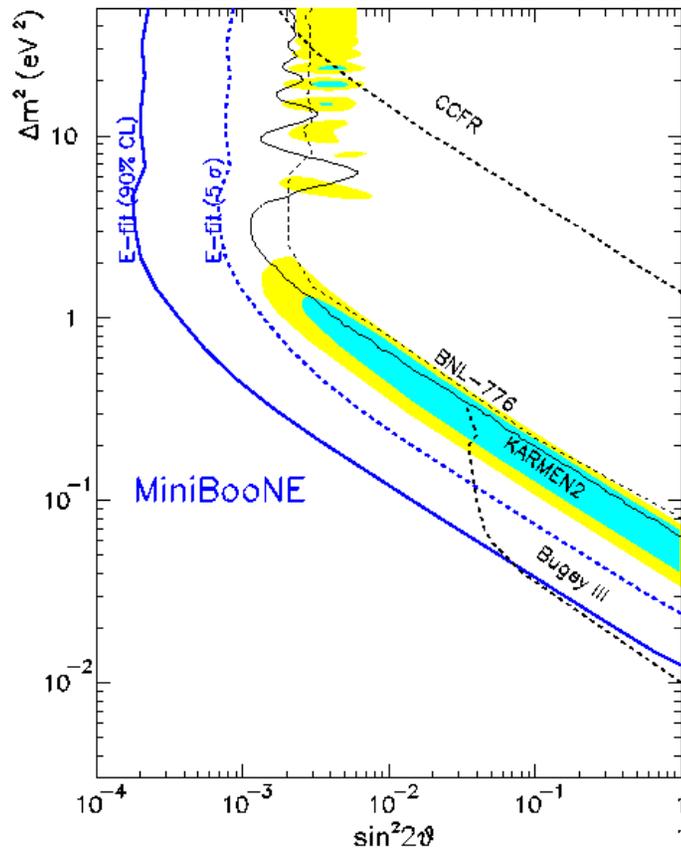
Use protons from
the 8 GeV booster
⇒ Neutrino Beam
 $\langle E_\nu \rangle \sim 1 \text{ GeV}$



12m sphere filled with
mineral oil and PMTs
located 500m from source



MiniBooNE Rates and Sensitivity



- Expected events
 - 500,000 ν_μ CC quasi-elastic
 - **~1000 n_e if LSND correct**
- Decisive investigation of LSND region
 - LSND $\rightarrow >5\sigma$ signal in MiniBooNE
 - Osc. signal has different energy than intrinsic ν_e
 - Experimental determinations of all backgrounds.
- If MiniBooNE sees a signal
 - \Rightarrow Very rich ν -osc. phenomenology (ν_{sterile} , ν_τ appearance, $\bar{\nu}\nu$

- If signal is observed in MiniBooNE, then add second detector at appropriate distance

P Two detector BooNE experiment

BooNE can measure:

$$\Delta m^2 \text{ to } \pm 0.014 \text{ eV}^2$$

$$\sin^2 2\theta \text{ to } \pm 0.002$$

MiniBooNE Detector and Beam Status



May 19

- About 50% of Photomultiplier tubes are installed
- Detector ready for commissioning in Dec. 01
- Beam preparations going well

Beginning of data taking - April, 2002



April 25



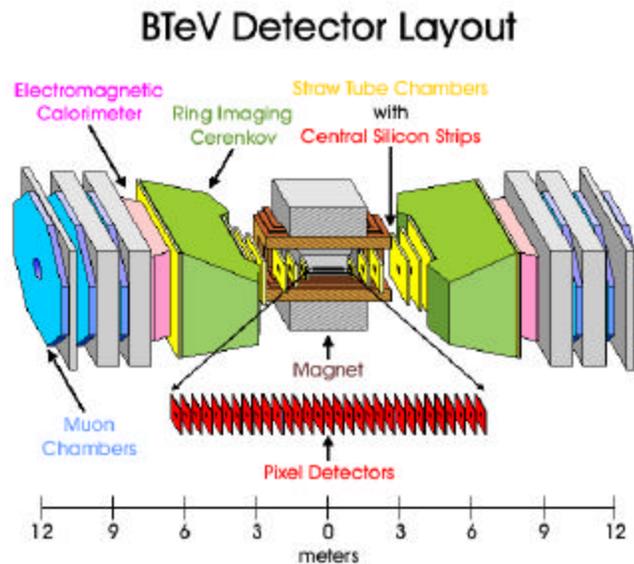
Quark Flavor Physics Studies at Fermilab

- Expect "New Physics" to show up as anomalous CP Violation
- Best measurements should have well understood, small experimental and theoretical errors.
 - $B_d \rightarrow \psi K_S, \rho \pi$
 - $B_s \rightarrow \psi \eta, D_s K$
 - $\Delta M_d / \Delta M_s$ in B_d and B_s decays
 - $K \rightarrow \pi \nu \bar{\nu}$
- CDF and D0
 - $B_d \rightarrow \psi K_S$ and $\sin(2\beta)$
 - x_s measurements
- BTeV
 - Wide range of B_s and B_d studies
 - $B_d \rightarrow \psi K_S, \rho \pi$; $B_s \rightarrow \psi \eta, D_s K$
- Rare Kaon Decays
 - CKM: $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ (SM: $\sim 8 \times 10^{-11}$)
 - KAMI: $K^0 \rightarrow \pi^0 \nu \bar{\nu}$ (SM: $\sim 3 \times 10^{-11}$)

Need to measure all of these processes since "New Physics" may change only one

BTeV: State-of-Art Hadronic B Experiment

- Physics
 - Many critical CP studies require the **large data samples** and **both B_d and B_s** available at hadron colliders
- BTeV Experiment
 - Innovative forward spectrometer PbWO₄ EM Calorimeter and RICH
 - Level 1 trigger based on pixels takes full advantage of enormous B cross section



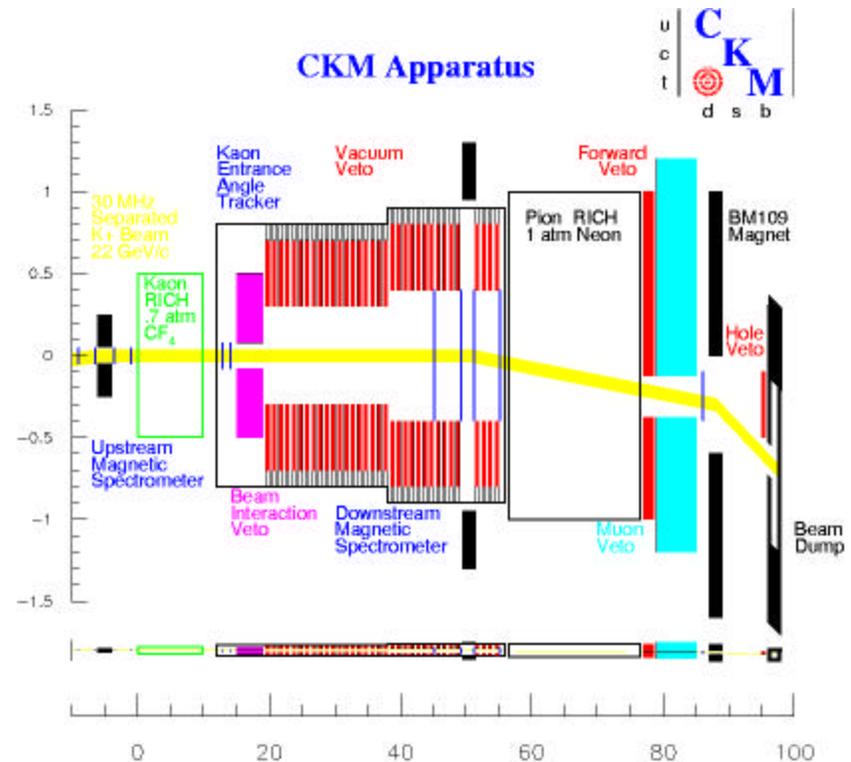
- Status:
 - Given Stage I approval after June, 00 Aspen PAC meeting based on:
 - Compelling physics that could be key in clarifying CP violation.
 - The BTeV detector unparalleled in its capabilities for this physics
- Next Steps and schedule:
 - BTeV doing R&D for Technical Design Report and working towards Stage II approval
 - Plan is for construction with a completion date in 2007

MI Kaon Program

- The rare processes $K \rightarrow \pi \nu \bar{\nu}$ are robust theoretically
 - Cleanliness comparable to the best B-meson time asymmetries
 - $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ measures V_{td} and $K^0 \rightarrow \pi^0 \nu \bar{\nu}$ measures η
- CKM and KAMI collaborations submitted proposals this Spring and Lab decision was made after recommendations from June PAC meeting
 - Key criteria was ability to measure BR at 10% level and credibly establish a discrepancy or agreement with B-physics parameters
- Lab decision:
 - CKM ($K^+ \rightarrow \pi^+ \nu \bar{\nu}$) was given Stage I approval
 - Convincing redundant experimental technique
 - Backgrounds can be understood from data
 - KAMI ($K^0 \rightarrow \pi^0 \nu \bar{\nu}$) was not given approval

CKM (Charged K⁺ Experiment)

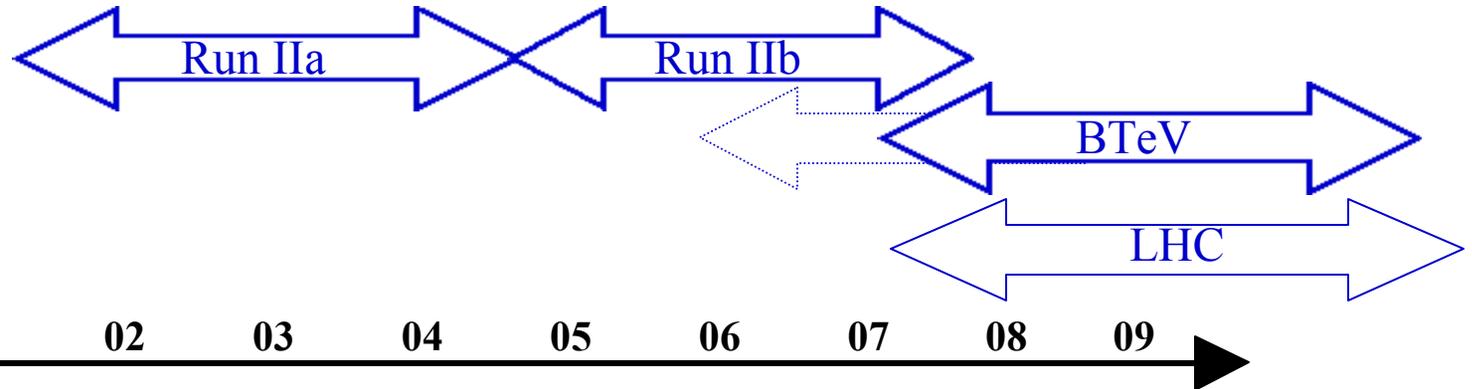
- Goal is to collect 100 K⁺ → π⁺νν events (Expect BR = 8±3×10⁻¹¹)
⇒ Measure |V_{td}| to 10%
(BNL787 has seen one event)
- Key component is 22 GeV super-conducting RF separated K⁺ beam
- Incident K⁺ and decay π⁺ velocities and momenta are measured with two high-speed redundant spectrometers.
- The two principle backgrounds of K⁺ → μ⁺ν and K⁺ → π⁺π⁰ are controlled with ultra-hermetic muon and photon veto systems
 - Bkgnd can be derived from obs. events



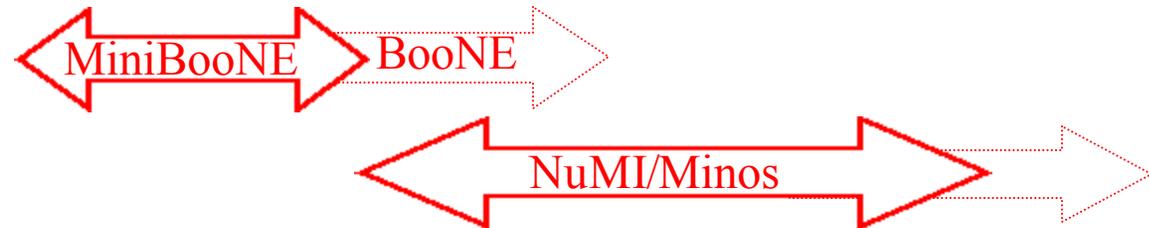
The CKM is a joint effort with BNL E-949 (follow-up to BNL E-787)

Fermilab Research Program

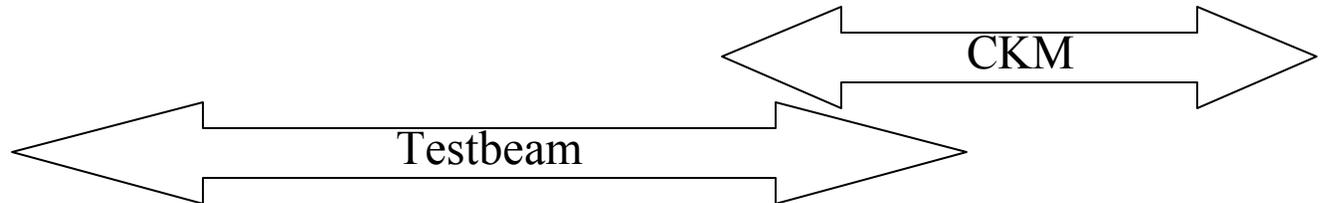
Collider:



Neutrinos:



MI Fixed Target:



Astrophysics:

